# How to determine the size of an asteroid ?

André Debackère, Monistrol sur Loire, FR <u>andre.debackere@free.fr</u> <u>http:dbrastronomie.free.fr</u>



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The method presented here is directly inspired by that described by Roger Dymock & Richard Miles.

#### List of steps

1) Preparation of the observation : choice of the target.

2) Acquiring images.

3) Downloading images.

4) Converting images into fits format suitable for ASTROMETRICA software.

5) Study of stars in the field of view aid of ALADIN software in order to find reference stars to apply photometry.

6) Vmag calculation for each reference star from r', J and Kmag.<sup>1</sup>

7) Target identification in the field.

8) Photometry of reference stars and target.

9) Vmag calculation for the target.

10) Absolute magnitude H calculation for the target.<sup>2</sup>

11) Diameter calculation for the target.

<sup>&</sup>lt;sup>1</sup> http://www.britastro.org/asteroids/JBAA 119 149-156 Dymock1.pdf

<sup>&</sup>lt;sup>2</sup> http://www.britastro.org/asteroids/dymock4.pdf

# Chapter I

# *I*) Choice of the target

1<sup>st</sup> criteria: the asteroid must be near opposition.

« Opposition » means that sun, earth, target are aligned so we can calculate the coordinates of the asteroid from the coordinates of the sun at the date of observation. You can use Gilbert JAVAUX website<sup>3</sup>

 $2^{nd}$  criteria : the magnitude of the asteroid must be below a value which depend of your device.

 $3^{rd}$  criteria : the phasis angle of the asteroid must be less than 1°.

 $4^{th}$  criteria : the altitude of the asteroid must be higher than about  $25^{\circ}/30^{\circ}$ .

5<sup>th</sup> criteria : the asteroid must be out of Earth umbra or penumbra

Now go to the Minor Planet Center website<sup>4</sup>

« OBSERVERS »
« Other Obervers Services »
« MPChecker »
This tool gives us a list of asteroids near opposition on the date of observation.

Then the Minor Planet Center Ephemeris Services gives us the parameters of the selected asteroid.

- « OBSERVERS »
- « Ephemeris Service »
- « designation or name »
- « Ephemeris start date »
- « Number of dates to output »
- « Ephemeris interval »
- « Observatory code »

You get a list from which you can choose the time of observation that gives you position and Vmag of the asteroid.

<sup>&</sup>lt;sup>3</sup> http://pgj.pagesperso-orange.fr/position-planetes.htm

<sup>&</sup>lt;sup>4</sup> http://www.minorplanetcenter.net/iau/mpc.html

### *II)* Images

**1**<sup>st</sup> search for good exposure time

At the bottom of the LCOGT home page click on « Instruments » then on « exposure\_time\_calculator ».<sup>5</sup>

Fill the form

- « Magnitude » Vmag given by MPC
- « ExpTime (sec) » try different value to avoid saturation
- « Telescope class and detector » T1m SBIG

« Filter » V

« Moon phase »

« Airmass » 1.3 by default

Click on « Calculate », then Calculated Values are updated :

- S/N:
- Magnitude:
- ExpTime(sec):
- PkDN:

2<sup>nd</sup> Request observations

Log into your LCOGT account

At the top of the LCOGT home page click on « Observatory »

« Schedule Observations > »

#### Observing

- « Request observations > »
- « Enter request title »
- « Queue schedule within »
- « Add an observation »
  - Add a target
  - Target name
  - RA
  - DEC
  - Use single filter ?
  - Exposure time ans How many times ?
  - Add
- « Submit »

*3<sup>rd</sup>* Keep fingers crossed...

<sup>&</sup>lt;sup>5</sup> http://lcogt.net/files/etc/exposure\_time\_calculator.html

#### *III)* Your images are ready

At the top of the LCOGT home page click on

« Observatory »

« Schedule Observations > »

#### Observations

« View submitted observations »

You can download your images

#### *IV)* Convert your images to be used in Astrometrica

If we use LCOGT network 1m telescopes equipped with SBIG camera we will convert the images obtained.

We use SalsaJ<sup>6</sup> software to perform this conversion.

The first thing to do after downloading the images on the LCOGT website is to open image with SalsaJ and save it in «FITS...» format, otherwise the image can not be used with Astrometrica.

Do it for each image you have downloaded.

#### *V)* Reference stars

We use ALADIN Sky Atlas<sup>7</sup> which allow us to display and align our own image, a Digital Sky Survey (DSS) image and the CMC14 catalogue.

- Edition « Select all objects »

- Copy all measures (to Excel).

- Open an Excel spreadsheet and then paste the list.

- Now looking for possible variable stars in the FOV using the catalogue : I / 280B.

**Criteria** for selection of stars from the previous list :

- r 'mag <~ 16
- 0.3 <J-K <0.7
- non-variable stars
- stars should not be saturated

<sup>&</sup>lt;sup>6</sup> http://www.fr.euhou.net/

<sup>&</sup>lt;sup>7</sup> http://aladin.u-strasbg.fr/

To measure the light intensity of each star we use Astrometrica<sup>8</sup>. First fill the different form in « Edit Program Settings »

- Observing site
- CCD
- Program
- Environnment
- Catalogs
- Internet

Then you download your images and click on « Astrometry », « Data Reduction… »

Click on a selected star, a window opens « Object Verification » and you can see SNR, Flux, FWHM. Do that for all the stars of the previous list.

Now you can choose the reference stars you want to use and you can calculate the V-band magnitude from r ', J and K using the formula :

# V = 0.6278 x (J-K) + 0.9947 x r '

# *VI)* Asteroid identification in the FOV

Click on « Tools », « Known Object Overlay »

Asteroids present in the FOV are marked in red.

Click on the asteroid you have chosen to study, you can see SNR, Flux and FWHM.

# *VII)* Image optimization

If the SNR of your images is smaller than 20, you can stack several images to increase the SNR but to avoid trailing when stacking images you should not exceed a maximum time value given by

I = FWHM/sky motion (Stephen Laurie law) with

- I in minutes

- FWHM in arcseconds

- sky motion in arcseconds per minute.

You can stack your images with Astrometrica according the number above.

- « Astrometry »
- « Track & Stack... »
- « Data Reduction... »
- « Tools »
- « Known Object Overlay »

- Click on the reference stars and on the asteroid to know the flux of each object.

<sup>&</sup>lt;sup>8</sup> http://www.astrometrica.at/

# *VIII)* V mag, H mag and "diameter" calculation

Now you can calculte the V-band magnitude of the asteroid by using LCOGT spreasheet<sup>9</sup>. I modified this spreadsheet to fit our work as you can see here: <u>https://onedrive.live.com/redir?resid=62E512265D1AC767!1331&authkey=!</u> <u>AEhgaszEbSYzAfQ&ithint=file%2cxlsx</u>

You need this information to complete it:

- Reference stars magnitude (§ V)
- Images exposure time (§ II)
- Reference stars flux (§ VII)

- Use the formula to calculate H mag:

#### $H = H(Ph) + 2.5log[(1-G)\phi_1(Ph) + G\phi_2(Ph)]$

where:

 $\varphi_i(Ph) = \exp\{-A_i(\tan 1/2 Ph)B_i\}$ i = 1 or 2,  $A_1 = 3.33$ ,  $A_2 = 1.87$ ,  $B_1 = 0.63$  and  $B_2 = 1.22$ and Ph is the phase angle in degrees given by the MPC (§ I). with  $H(Ph) = V - 5 \log(r \cdot Delta)$  r and Delta given by the MPC (§ I)

- Choose the geometric albedo pV

There are many types of asteroids, but the most common are:

- type C (75%) -> pV=0.4
- types S (17%) and M -> pV=0.15

# $D(km) = 1329 \times 10^{(-H/5)} / pV^{1/2}$

Note that the diameter we get is an "equivalent photometric diameter" since it is unlikely that the asteroids studied are perfectly spherical.

<sup>&</sup>lt;sup>9</sup> http://resources.faulkes-telescope.com/course/view.php?id=39

#### I) Observation date: 20141026

- Coordinates of the Sun (PGJ, chapter I, § I, 1<sup>st</sup> criteria) 14h01m13.34s -12°20'16.83"



- Opposition
   RA= 14h01m13.34s + 12h = 02h01m13.34s
   DEC = +12°20'16.83"
- MPChecker (chapter I, § I, 5<sup>th</sup> criteria)
   Vmag limit = 18.0
   60.0 arcminutes

Produce list Clear/reset form		
Date : 2014 10 26.0 🗸 UT		
Produce list of known minor planets around:		
It is J2000.0 position: R.A. = 02 01 13.34 Decl. = +12 20 16.83		
or around 🔘 these observations:		
.::		
Radius of search = 60 arc-minutes		
Limiting magnitude, V = 18.0 Observatory code = 500		
Output matches in order of:		
increasing distance from specified position		
Display motions in arcseconds per 🔘 minute or 💿 hour		
or 🕜 degrees per day		
Disptay () total or () separate motions		
Output designations in 💿 unpacked or 🔘 packed form		
Output:		
all objects		

MPChecker generates a list of asteroids near opposition Choice : **(48644) 1995 UG7** 

MPC OBSERVERS Ephemeris Service (chapter I, § I, 5<sup>th</sup> criteria) designation : 48644
 Ephemeris start date : 20141026
 Number of dates to output : 144
 Ephemeris interval : 10 minutes
 Observatory code : 500 (center of the Earth by default)

Then click on « Get ephemerides » Note the different parameters

#### 11h20 UT

- R.A. (J2000) = 02 :01 :32.5 Decl. (J2000) = +11 :55 :44 Delta = 1.352 AU r = 2.346 AU El. = 179.5° Ph. = 0.2° V = 17.1Sky Motion = 0.67 "/min Alt. = +34°
- **Request observation settings** -« Enter request title » : aster size « Queue schedule within » : 24 hours « What size telescope ? » : 1-meter

  - « Target name » : 48644
  - « RA » : 02 01 35.5
  - « DEC » :+11 57 50.0
  - « Filters » : Bessell-V
  - « Exposure time » : 40 s
  - « How many times ? » : 14



#### Aladin and CMC-14 catalog

There are 4 stars that meet the selection criteria.

1- 020121.0+120045	$\rightarrow$ Vmag = 14.63
2- 020139.6+120032	$\rightarrow$ Vmag = 15.27
3- 020144.4+115902	$\rightarrow$ Vmag = 15.15
4- 020124.9+115509	$\rightarrow$ Vmag = 15.10





Improvement of S/N

- FWHM average 1.74"
- max time interval 2.59 minutes
- full duration 12.7 minutes
- nb images 4 (4x3 stacked and 2 rejected)

Star 1 is saturated on 2 images -> calculation with 3 reference stars (stars 2,3 &4) Average Vmag for the asteroid : 17.214 Hmag : 15.4531131

With albedo 0.4 (type C) **D = 1.7 km** 

With albedo 0.15 (type S or M) **D = 2.8 km** 

#### II) Observation date: 20141111

PGJ website : enter the date: Day 11 ; Month 11 ; Year 2014, you can see the coordinates of the sun RASun = 15h07m13.41s and DECSun=  $-17^{\circ}32'13.55''$ 

Then the coordinates of the targ	jet are :
RATarget = RASun + 12h	RATarget = 03h07m13.41s
DECTarget = opp(DECSun)	DECTarget = +17°32'13.55''

Now, open MPChecker a tool of the Minor Planet Center

« OBSERVERS » « Other Obervers Services » « MPChecker » then fill the form : Date : year month day (UT) 2014 11 11.50 means 2014 November 11th at 12h UT Produce the list of known minor planets around : this J2000.0 position: R.A. = 03 07 13.41 Decl. = +17 32 13.55 Radius of search in arc-minutes 120 Limiting magnitude, V = 16 Observatory code = 500 Click on « Produce list » we get 6 asteroids You can modify the radius of search (decrease for less asteroids), with a radius of 60 arc-minutes we get only one asteroid : (12735) 1991 VV1 Now, open the Minor Planet Center webpage « OBSERVERS » « Ephemeris Service » and fill the form Objects may be identified by designation or by name. Enter a list of designations or names below (one entry per line, excess entries will be ignored): 12735 Ephemeris start date: 20141111 Number of dates to output 144 **Ephemeris** interval:

10 minutes

Observatory code : V37 (we choose to observe from Mac Donald Observatory)

then click on « Get ephemerides/HTML page » this is the result : Date : 2014 11 11 UT 120000 R.A (J2000) 03 05 45.1 Decl. +17 27 33 Delta : 1.066 r : 2.057 El. : 179.9 Ph : 0.0 V :15.4 Sky motion : 0.78 "/min Object Alt. : +17 ....... In Earth's penumbra

It's not a good choice !!!